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# Geometry of Admissible Curves of Constant-Ratio in Pseudo-Galilean Space 

M. Khalifa Saad ${ }^{1, *}$, H. S. Abdel-Aziz ${ }^{2}$, Haytham A. Ali $^{2}$<br>${ }^{1}$ Department of Mathematics, Faculty of Science, Islamic University of Madinah, KSA<br>${ }^{2}$ Department of Mathematics, Faculty of Science, Sohag University, 82524 Sohag, Egypt<br>*Corresponding author: mohammed.khalifa@iu.edu.sa


#### Abstract

An admissible curve of a pseudo-Galilean space is said to be of constant-ratio if the ratio of the length of the tangent and normal components of its position vector function is a constant. In this paper, we investigate and characterize a spacelike admissible curve of constant-ratio in terms of its curvature functions in the pseudo-Galilean space $G_{3}^{1}$. Also, we study some special curves of constantratio such as $T$-constant and $N$-constant types of these curves. Finally, we give some computational examples for constructing the meant curves to demonstrate our theoretical results.


## 1. Introduction

According to the space curve theory, it is well known that, a curve $\alpha(s)$ in $E^{3}$ lies on a sphere if its position vector lies on its normal plane at each point. If the position vector $\alpha$ lies on its rectifying plane then $\alpha(s)$ is called a rectifying curve [1]. Rectifying curves are characterized by the simple equation:

$$
\begin{equation*}
\alpha(s)=\lambda(s) T(s)+\mu(s) B(s), \tag{1.1}
\end{equation*}
$$

where $\lambda(s)$ and $\mu(s)$ are smooth functions and $T(s)$ and $B(s)$ are tangent and binormal vector fields of $\alpha$, respectively. In [2] the author provided that a twisted curve is congruent to a non constant linear function of $s$. On the other hand, in the Minkowski 3-space $E_{1}^{3}$, the rectifying curves were investigated in $[3,4]$. Besides, in [4] a characterization of the spacelike, the timelike and the null rectifying curves in the Minkowski 3-space in terms of centrodes were given. The characterization of rectifying curves in three dimensional compact Lee groups as well as in dual spaces were given in [5], [6], respectively. For the study of constant-ratio curves, the authors gave the necessary and sufficient conditions for

[^0]curves in Euclidean and Minkowski spaces to become $T$-constant or $N$-constant [7-10]. In analogy with the Euclidean 3-dimensional case, our main goal in this work is to define the spacelike admissible curves of constant-ratio in the pseudo Galilean 3-space as a curve whose position vector always lies in the orthogonal complement $N^{\perp}$ of its principal normal vector field $N$. Consequently, $N^{\perp}$ is given by
$$
N^{\perp}=\left\{V \in G_{3}^{1}:<V, N>=0\right\},
$$
where $\langle\cdot, \cdot\rangle$ denotes the inner product in $G_{3}^{1}$. Hence $N^{\perp}$ is a 2-dimensional plane of $G_{3}^{1}$, spanned by the tangent and binormal vector fields $T$ and $B$, respectively. Therefore, the position vector with respect to some chosen origin of a considered curve $\alpha$ in $G_{3}^{1}$, satisfies the parametric equation:
\[

$$
\begin{equation*}
\alpha(s)=m_{o}(s) T(s)+m_{1}(s) N(s)+m_{2}(s) B(s), \tag{1.2}
\end{equation*}
$$

\]

for some differential functions $m_{i}(s), 0 \leq i \leq 2$, where $s$ is arc-length parameter. Then, we give the necessary and sufficient conditions for the curve $\alpha$ in $G_{3}^{1}$ to be a constant-ratio curve.

## 2. Pseudo-Galilean geometry

In this section, we introduce the basic concepts, familiar definitions and notations on pseudoGalilean space which are needed throughout this study. The pseudo-Galilean geometry is one of the real Cayley-Klein geometries of projective signature ( $0,0,+,-$ ). The absolute of the pseudo-Galilean geometry is an ordered triple $\{w, f, I\}$ where $w$ is the ideal (absolute) plane, $f$ is a line in $w$ and $I$ is the fixed hyperbolic involution of points of $f$, for more details, we refer to [11,12]. The geometry of the pseudo-Galilean space is similar (but not the same) to the Galilean space which was presented in [11]. The inner and cross product of two vectors $\mathbf{x}=\left(x_{1}, y_{1}, z_{1}\right)$ and $\mathbf{y}=\left(x_{2}, y_{2}, z_{2}\right)$ in $G_{3}^{1}$ are, respectively defined as follows:

$$
\begin{gathered}
g(\mathbf{x}, \mathbf{y})= \begin{cases}x_{1} x_{2}, & \text { if } x_{1} \neq 0 \vee x_{2} \neq 0, \\
y_{1} y_{2}-z_{1} z_{2} & \text { if } x_{1}=0 \wedge x_{2}=0,\end{cases} \\
\mathbf{x} \times \mathbf{y}=\left|\begin{array}{ccc}
0 & -e_{2} & e_{3} \\
x_{1} & y_{1} & z_{1} \\
x_{2} & y_{2} & z_{2}
\end{array}\right| .
\end{gathered}
$$

Also the norm of a vector $\mathbf{x}=(x, y, z)$ is given by

$$
\|\mathbf{x}\|=\left\{\begin{array}{cl}
x & , \text { if } x \neq 0  \tag{2.1}\\
\sqrt{\left|y^{2}-z^{2}\right|} & , \text { if } x=0
\end{array}\right.
$$

The group of motions of the pseudo-Galilean $G_{3}^{1}$ is a six-parameter group given (in affine coordinates) by

$$
\begin{aligned}
& \bar{x}=a+x \\
& \bar{y}=b+c x+y \cosh \varphi+z \sinh \varphi \\
& \bar{z}=d+e x+y \sinh \varphi+z \cosh \varphi
\end{aligned}
$$

According to the motion group in pseudo-Galilean space, a vector $\mathbf{x}(x, y, z)$ is said to be non isotropic if $x \neq 0$. All unit non-isotropic vectors are of the form $(1, y, z)$. For isotropic vectors, $x=0$ holds. There are four types of isotropic vectors: spacelike $\left(y^{2}-z^{2}\right)>0$, timelike $\left(y^{2}-z^{2}\right)<0$, and two types of lightlike $(y= \pm z)$ vectors. A non-lightlike isotropic vector is a unit vector if $y^{2}-z^{2}= \pm 1$.

A trihedron $\left(T_{0} ; e_{1}, e_{2}, e_{3}\right)$ with a proper origin $T_{0}\left(x_{0}, y_{0}, z_{0}\right)$ which is orthonormal in pseudoGalilean sense if the vectors $e_{1}, e_{2}, e_{3}$ are of the following form: $e_{1}=\left(1, y_{1}, z_{1}\right), e_{2}=\left(0, y_{2}, z_{2}\right)$ and $e_{3}=\left(0, \varepsilon z_{2}, \varepsilon y_{2}\right)$ with $y^{2}-z^{2}=\delta$, where $\varepsilon, \delta$ is +1 or -1 . Such trihedron $\left(T_{o} ; e_{1}, e_{2}, e_{3}\right)$ is called positively oriented if for its vectors, $\operatorname{det}\left(e_{1}, e_{2}, e_{3}\right)=1$ holds; that is if $y^{2}-z^{2}=\varepsilon$.

Let $\alpha(t): I \subset R \rightarrow G_{3}^{1}$ be a curve parameterized by $\alpha(t)=(x(t), y(t), z(t))$, where $x(t), y(t), z(t) \in C^{3}$ (the set of three-times continuously differentiable functions) and $t$ run through a real interval [12].

Definition 2.1. A curve $\alpha$ given by $\alpha(t)=(x(t), y(t), z(t))$ is admissible if $\dot{x}(t) \neq 0$.
Also, If $\alpha$ is taken as follows:

$$
\begin{equation*}
\alpha(x)=(x, y(x), z(x)) \tag{2.2}
\end{equation*}
$$

with the condition

$$
\begin{equation*}
y^{\prime \prime 2}(x)-z^{\prime \prime 2}(x) \neq 0 \tag{2.3}
\end{equation*}
$$

then the arc-length parameter $s$ is defined by

$$
\begin{equation*}
d s=|\dot{x}(t) d t|=d x \tag{2.4}
\end{equation*}
$$

Here, we assume that $d s=d x$ and $s=x$ as the arc-length of the curve $\alpha$ [12]. The vector

$$
T(s)=\alpha^{\prime}(s)
$$

is called the tangent unit vector of $\alpha$. Also, the unit vector field is given by

$$
\begin{equation*}
N(s)=\frac{\alpha^{\prime \prime}(s)}{\sqrt{\left|y^{\prime \prime 2}(s)-z^{\prime \prime 2}(s)\right|}} \tag{2.5}
\end{equation*}
$$

and the binormal vector is expressed as

$$
\begin{equation*}
B(s)=\frac{\left(0, \varepsilon z^{\prime \prime}(s), \varepsilon y^{\prime \prime}(s)\right)}{\sqrt{\left|y^{\prime \prime 2}(s)-z^{\prime \prime 2}(s)\right|}} \tag{2.6}
\end{equation*}
$$

and it is orthogonal in pseudo-Galilean sense to the osculating plane of $\alpha$ spanned by the vectors $\alpha^{\prime}(s)$ and $\alpha^{\prime \prime}(s)$. The curve $\alpha$ given by Eq. (2.2) is a spacelike (resp. timelike) if $N(s)$ is a timelike (resp.
spacelike) vector. The principal normal vector or simply normal is spacelike if $\varepsilon=+1$ and timelike if $\varepsilon=-1$. Here $\varepsilon=+1$ or -1 is chosen by the criterion $\operatorname{det}(T, N, B)=1$. That means

$$
\begin{equation*}
\left|y^{\prime \prime 2}(s)-z^{\prime \prime 2}(s)\right|=\varepsilon\left(y^{\prime \prime 2}(s)-z^{\prime \prime 2}(s)\right) \tag{2.7}
\end{equation*}
$$

Definition 2.2. In each point of an admissible curve in $G_{3}^{1}$, the associated orthonormal (in pseudoGalilean sense) trihedron $\{T(s), N(s), B(s)\}$ can be defined. This trihedron is called pseudo-Galilean Frenet trihedron.

For the pseudo-Galilean Frenet trihedron of an admissible curve $\alpha$, the Frenet equations are defined as:

$$
\begin{align*}
T^{\prime} & =\kappa N \\
N^{\prime} & =\tau B  \tag{2.8}\\
B^{\prime} & =\tau N
\end{align*}
$$

where $\kappa$ and $\tau$ are the pseudo-Galilean curvatures of $\alpha$ defined as follows:

$$
\begin{gather*}
\kappa(s)=\sqrt{\left|y^{\prime \prime 2}(s)-z^{\prime \prime 2}(s)\right|}  \tag{2.9}\\
\tau(s)=\frac{y^{\prime \prime}(s) z^{\prime \prime \prime}(s)-y^{\prime \prime \prime}(s) z^{\prime \prime}(s)}{\kappa^{2}(s)} \tag{2.10}
\end{gather*}
$$

and the pseudo-Galilean torsion can be written in the form

$$
\begin{equation*}
\tau(s)=\frac{\operatorname{det}\left(\alpha^{\prime}(s), \alpha^{\prime \prime}(s), \alpha^{\prime \prime \prime}(s)\right)}{\kappa^{2}(s)} \tag{2.11}
\end{equation*}
$$

The Serret-Frenet equations (2.8) can be written in matrix form as

$$
\frac{d}{d s}\left[\begin{array}{c}
T \\
N \\
B
\end{array}\right]=\left[\begin{array}{ccc}
0 & \kappa & 0 \\
0 & 0 & \tau \\
0 & \tau & 0
\end{array}\right]\left[\begin{array}{c}
T \\
N \\
B
\end{array}\right]
$$

The Pseudo-Galilean sphere with radius $r$ is defined by

$$
S_{ \pm}^{2}=\left\{u \in G_{3}^{1}: g(u, u)= \pm r^{2}\right\}
$$

3. Spacelike curves of constant-ratio in $G_{3}^{1}$

Let $\alpha: I \subset R \rightarrow G_{3}^{1}$ be an arbitrary spacelike admissible curve. In the light of which introduced in [13-15], we consider the following theorem.

Theorem 3.1. The position vector of $\alpha$ with curvatures $\kappa(s)$ and $\tau(s) \neq 0$, and with respect to the Frenet frame in the pseudo-Galilean space $G_{3}^{1}$, it can be written as

$$
\begin{align*}
\alpha= & \left(s+c_{0}\right) T+e^{-\int \tau(s) d s}\left(c_{1} e^{2 \int \tau(s) d s}+e^{2 \int \tau(s) d s} \int \frac{\kappa(s)\left(s+c_{0}\right)}{2} e^{-\int \tau(s) d s} d s\right. \\
& \left.-\int \frac{\kappa(s)\left(s+c_{0}\right)}{2} e^{\int \tau(s) d s} d s+c_{2}\right) N+e^{-\int \tau(s) d s}\left(c_{1} e^{2 \int \tau(s) d s}\right.  \tag{3.1}\\
& \left.+e^{2 \int \tau(s) d s} \int \frac{\kappa(s)\left(s+c_{0}\right)}{2} e^{-\int \tau(s) d s} d s+\int \frac{\kappa(s)\left(s+c_{0}\right)}{2} e^{\int \tau(s) d s} d s-c_{2}\right) B .
\end{align*}
$$

where $c_{0}, c_{1}$ and $c_{2}$ are arbitrary constants.
Proof. Let $\alpha$ be an arbitrary spacelike curve in the pseudo-Galilean space $G_{3}^{1}$, then we may express its position vector as

$$
\alpha(s)=m_{0}(s) T(s)+m_{1}(s) N(s)+m_{2}(s) B(s) .
$$

Differentiating this equation with respect to the arc-length parameter $s$ and using the Serret-Frenet equations (2.8), we obtain

$$
\begin{aligned}
\alpha^{\prime}(s)= & m_{o}^{\prime}(s) T(s)+\left(m_{1}^{\prime}(s)+\kappa(s) m_{o}(s)+\tau(s) m_{2}(s)\right) N(s) \\
& +\left(m_{2}^{\prime}(s)+\tau(s) m_{1}(s)\right) B(s),
\end{aligned}
$$

it follows that

$$
\begin{align*}
m_{o}^{\prime}(s) & =1, \\
m_{1}^{\prime}(s)+\kappa(s) m_{0}(s)+\tau(s) m_{2}(s) & =0,  \tag{3.2}\\
m_{2}^{\prime}(s)+\tau(s) m_{1}(s) & =0 .
\end{align*}
$$

From Eqs. (3.2), we have

$$
\begin{equation*}
m_{0}(s)=s+c_{0} . \tag{3.3}
\end{equation*}
$$

It is useful to change the variable $s$ to the variable $t=\int \tau(s) d s$. Therefore all functions of $s$ will transform to the functions of $t$. Here, we will use dot to denote the derivative with respect to $t$ (where the prime denotes the derivative with respect to $s$ ). Also, From Eq. (3.2), we get

$$
\begin{equation*}
m_{1}(t)=-\dot{m}_{2}(t), \quad \text { where } \dot{m}_{2}=\frac{d m_{2}}{d t} \tag{3.4}
\end{equation*}
$$

it leads to

$$
\begin{equation*}
\ddot{m}_{2}(t)-m_{2}(t)=\frac{y(t) \kappa(t)}{\tau(t)}, \quad y(t)=m_{0}(s)=s+c_{0} . \tag{3.5}
\end{equation*}
$$

The general solution of this equation is given by

$$
\begin{equation*}
m_{2}(t)=e^{-t}\left[c_{1} e^{2 t}+e^{2 t} \int \frac{\kappa(t) y(t)}{2 \tau(t)} e^{-t} d t+\int \frac{\kappa(t) y(t)}{2 \tau(t)} e^{t} d t-c_{2}\right], \tag{3.6}
\end{equation*}
$$

where $c_{1}$ and $c_{2}$ are arbitrary constants. From Eqs. (3.4) and (3.6), we obtain the function $m_{1}(t)$ as

$$
\begin{equation*}
m_{1}(t)=e^{-t}\left[c_{1} e^{2 t}+e^{2 t} \int \frac{\kappa(t) y(t)}{2 \tau(t)} e^{-t} d t-\int \frac{\kappa(t) y(t)}{2 \tau(t)} e^{t} d t+c_{2}\right] . \tag{3.7}
\end{equation*}
$$

Hence, Eqs. (3.6) and (3.7) take the following forms:

$$
\begin{align*}
& m_{1}=e^{-\int \tau(s) d s}\left[c_{1} e^{2 \int \tau(s) d s}+e^{2 \int \tau(s) d s} \int \frac{\left(s+c_{0}\right) \kappa}{2} e^{-\int \tau(s) d s} d s-\int \frac{\left(s+c_{0}\right) \kappa}{2} e^{\int \tau(s) d s} d s+c_{2}\right],  \tag{3.8}\\
& m_{2}=e^{-\int \tau(s) d s}\left[c_{1} e^{2 \int \tau(s) d s}+e^{2 \int \tau(s) d s} \int \frac{\left(s+c_{0}\right) \kappa}{2} e^{-\int \tau(s) d s} d s+\int \frac{\left(s+c_{0}\right) \kappa}{2} e^{\int \tau(s) d s} d s-c_{2}\right] . \tag{3.9}
\end{align*}
$$

Substituting from Eqs. (3.3), (3.8) and (3.9) in Eq. (1.2), the result (3.1) is obtained and thus, the proof is completed.

Theorem 3.2. Let $\alpha: I \subset R \rightarrow G_{3}^{1}$ be a spacelike curve with $\kappa \neq 0$ and $\tau \neq 0$ in $G_{3}^{1}$. Then the position vector and curvatures of $\alpha$ satisfy a vector differential equation of third order.

Proof. Let $\alpha: I \subset R \rightarrow G_{3}^{1}$ be a spacelike curve with curvatures $\kappa \neq 0$ and $\tau \neq 0$ in $G_{3}^{1}$. From Frenet equations (2.8), one can write

$$
\begin{align*}
N & =\frac{T^{\prime}}{\kappa}  \tag{3.10}\\
B & =\frac{N^{\prime}}{\tau} . \tag{3.11}
\end{align*}
$$

Substituting Eq. (3.10) in Eq. (2.8), we get

$$
\begin{equation*}
B^{\prime}=\frac{\tau}{\kappa} T^{\prime} \tag{3.12}
\end{equation*}
$$

Differentiating Eq. (3.10) with respect to $s$ and substituting in Eq. (3.10), we find

$$
\begin{equation*}
B=\frac{1}{\tau}\left[\left(\frac{1}{\kappa}\right)^{\prime} T^{\prime}+\left(\frac{1}{\kappa}\right) T^{\prime \prime}\right] \tag{3.13}
\end{equation*}
$$

Similarly, taking the differentiation of Eq. (3.13) and equalize with Eq. (2.8), we obtain

$$
\begin{equation*}
\frac{1}{\tau \kappa} T^{\prime \prime \prime}+\left[2 \frac{1}{\tau}\left(\frac{1}{\kappa}\right)^{\prime}-\left(\frac{1}{\tau}\right)^{\prime} \frac{1}{\kappa}\right] T^{\prime \prime}+\left[\frac{1}{\tau}\left(\left(\frac{1}{\kappa}\right)^{\prime \prime}-\frac{\tau^{2}}{\kappa}\right)-\left(\frac{1}{\tau}\right)^{\prime}\left(\frac{1}{\kappa}\right)^{\prime}\right] T^{\prime}=0 \tag{3.14}
\end{equation*}
$$

Hence, it completes the proof.
Theorem 3.3. The position vector $\alpha(s)$ of a spacelike admissible curve with curvature $\kappa(s)$ and torsion $\tau(s)$ in the pseudo-Galilean space $G_{3}^{1}$ is computed from the intrinsic representation form

$$
\alpha(s)=\left(s,-\int\left[\int \kappa(s) \sinh \left[\int \tau(s) d s\right] d s\right] d s, \int\left[\int \kappa(s) \cosh \left[\int \tau(s) d s\right] d s\right] d s\right)
$$

with tangent, principal normal and binormal vectors respectively, are given by

$$
\begin{aligned}
& T(s)=\left(1,-\int \kappa(s) \sinh \left[\int \tau(s) d s\right] d s, \int \kappa(s) \cosh \left[\int \tau(s) d s\right] d s\right) \\
& N(s)=\left(0,-\sinh \left[\int \tau(s) d s\right], \cosh \left[\int \tau(s) d s\right]\right) \\
& B(s)=\left(0,-\cosh \left[\int \tau(s) d s\right], \sinh \left[\int \tau(s) d s\right]\right) .
\end{aligned}
$$

Now, for each given $\alpha: I \subset R \rightarrow G_{3}^{1}$, there is a natural orthogonal decomposition of the position vector $\alpha$ at each point on $\alpha$; namely,

$$
\begin{equation*}
\alpha=\alpha^{T}+\alpha^{N} \tag{3.15}
\end{equation*}
$$

where $\alpha^{T}$ and $\alpha^{N}$ denote the tangential and normal components of $\alpha$ at the point, respectively. Let $\left\|\alpha^{T}\right\|$ and $\left\|\alpha^{N}\right\|$ denote the length of $\alpha^{T}$ and $\alpha^{N}$, respectively. In what follows we introduce the notion of constant-ratio curves. So, similar to the Euclidean case [16], we consider the following definitions [17].

Definition 3.1. A curve $\alpha$ of the pseudo-Galilean space $G_{3}^{1}$ is said to be of constant-ratio curve if the ratio $\left\|\alpha^{T}\right\|:\left\|\alpha^{N}\right\|$ is constant on $\alpha(I)$.

Clearly, for a constant-ratio curve in $G_{3}^{1}$, we have

$$
\begin{equation*}
\frac{m_{o}^{2}}{m_{2}^{2}-m_{1}^{2}}=c_{3} \tag{3.16}
\end{equation*}
$$

for some constant $c_{3}$.
Definition 3.2. Let $\alpha: I \subset R \rightarrow G_{3}^{1}$ be an admissible curve in $G_{3}^{1}$. If $\left\|\alpha^{T}\right\|$ is constant, then $\alpha$ is called $T$-constant curve. Further, $T$-constant curve $\alpha$ is called of first kind if $\left\|\alpha^{T}\right\|=0$, otherwise is called of second kind.

Definition 3.3. Let $\alpha: I \subset R \rightarrow G_{3}^{1}$ be an admissible curve in $G_{3}^{1}$. If $\left\|\alpha^{N}\right\|$ is constant, then $\alpha$ is called a $N$-constant curve. For a $N$-constant curve $\alpha$, either $\left\|\alpha^{N}\right\|=0$ or $\left\|\alpha^{N}\right\|=\mu$ for some nonzero smooth function $\mu$. Further, a $N$-constant curve $\alpha$ is called of first kind if $\left\|\alpha^{N}\right\|=0$, otherwise it is of second kind.

For $N$-constant curve $\alpha$ in $G \frac{1}{3}$, we can write

$$
\begin{equation*}
\left\|\alpha^{N}(s)\right\|^{2}=m_{2}^{2}(s)-m_{1}^{2}(s)=c_{4} \tag{3.17}
\end{equation*}
$$

where $c_{4}$ is constant.
In what follows, we characterize the admissible curves in terms of their curvature functions $m_{i}(s)$ and give the necessary and sufficient conditions for these curves to be $T$-constant or N -constant curves.

Theorem 3.4. Let $\alpha: I \subset R \rightarrow G_{3}^{1}$ be a spacelike curve in $G_{3}^{1}$. Then $\alpha$ is of constant-ratio if and only if

$$
\left(\frac{\kappa^{\prime}-\kappa^{3} c_{3}\left(s+c_{0}\right)}{c_{3} \kappa^{2} \tau}\right)^{\prime}=\frac{-\tau}{c_{3} \kappa} .
$$

Proof. Let $\alpha: I \subset R \rightarrow G_{3}^{1}$ be a spacelike curve given with the invariant parameter $s$. Then, we have

$$
m_{0}(s)=s+c_{0},
$$

where $c_{o}$ is an arbitrary constant. Also, from Eq. (3.16), the curvature functions $m_{i}(s), 0 \leq i \leq 2$ satisfy

$$
\begin{equation*}
m_{2}(s) m_{2}^{\prime}(s)-m_{1}(s) m_{1}^{\prime}(s)=\frac{s+c_{0}}{c_{3}} . \tag{3.18}
\end{equation*}
$$

By using Eqs. (3.2) with Eq. (3.18), we obtain

$$
m_{1}=\frac{1}{c_{3} k},
$$

it follows that

$$
m_{2}=\frac{\kappa^{\prime}-\kappa^{3} c_{3}\left(s+c_{0}\right)}{c_{3} \kappa^{2} \tau}
$$

thus, the result is clear.

## 3.1. $\mathbf{T}$-constant spacelike curves in $G_{3}^{1}$.

Proposition 3.1. There are no $T$-constant spacelike curves in pseudo-Galilean space $G_{3}^{1}$.
Proof. Let $\alpha: I \subset R \rightarrow G_{3}^{1}$ be a spacelike curve in $G_{3}^{1}$. Then $\left\|\alpha^{T}\right\|=m_{0}$, where $m_{0}$ is equal to zero or a nonzero constant. Since $m_{0}=x+c_{0}$, this contradicts the fact of value of $m_{0}$.
3.2. $N$-constant spacelike curves in $G_{3}^{1}$.

Lemma 3.1. Let $\alpha: I \subset R \rightarrow G_{3}^{1}$ be a spacelike curve in $G_{3}^{1}$. Then $\alpha$ is $N$-constant curve if and only if the following condition:

$$
m_{2}(s) m_{2}^{\prime}(s)-m_{1}(s) m_{1}^{\prime}(s)=0,
$$

holds together Eqs. (3.2), where $m_{i}(s), 0 \leq i \leq 2$ are differentiable functions.
Proposition 3.2. Let $\alpha: I \subset R \rightarrow G_{3}^{1}$ be a spacelike curve in $G_{3}^{1}$. Then $\alpha$ is a $N$-constant curve of first kind if $\alpha$ is a straight line in $G_{3}^{1}$.

Proof. Suppose that $\alpha$ is $N$-constant curve of first kind in $G_{3}^{1}$, then

$$
m_{2}^{2}(s)-m_{1}^{2}(s)=0
$$

So, we have two cases to be discussed:

## Case 1.

$$
m_{2}(s)=m_{1}(s)
$$

Using Eqs. (3.2), we get

$$
\kappa=0 .
$$

## Case 2.

$$
m_{2}(s)=-m_{1}(s) .
$$

Also, from Eqs. (3.2), we obtain

$$
\kappa=0 .
$$

It means that the curve $\alpha$ is a straight line in $G_{3}^{1}$.

Theorem 3.5. Let $\alpha: I \subset R \rightarrow G_{3}^{1}$ be a spacelike curve in $G_{3}^{1}$. If $\alpha$ is $N$-constant curve of second kind, then the position vector $\alpha$ has the parametrization:

$$
\begin{align*}
\alpha(s)= & \left(s+c_{0}\right) T(s)+\left[\frac{1}{4} e^{-u(s)}\left(-4 c_{4}+e^{2 u(s)}\right)-\frac{1}{2} e^{u(s)}\right] N(s) \\
& +\left[\frac{1}{4} e^{-u(s)}\left(-4 c_{4}+e^{2 u(s)}\right)\right] B(s), \tag{3.19}
\end{align*}
$$

where $u(s)=\int \tau(s) d s+c_{5}, c_{5}$ is integral constant.
Proof. From Eq. (3.3), we have

$$
m_{o}(s)=\left(s+c_{0}\right) .
$$

Besides, from of Eq. (3.2) and Eq. (3.17), we obtain

$$
m_{2}^{\prime 2}(s)-\tau^{2}(s) m_{2}^{2}(s)-c_{4} \tau^{2}(s)=0
$$

where $c_{4} \neq 0$ is a real constant. The solution of this equation is given by

$$
\begin{equation*}
m_{2}(s)=\frac{1}{4} e^{-u(s)}\left(-4 c_{4}+e^{2 u(s)}\right) . \tag{3.20}
\end{equation*}
$$

If we substitute Eq. (3.3) in Eq. (3.2), we can get

$$
\begin{equation*}
m_{1}(s)=\frac{1}{4} e^{-u(s)}\left(-4 c_{4}+e^{2 u(s)}\right)-\frac{1}{2} e^{u(s)}, \tag{3.21}
\end{equation*}
$$

hence, in light of Eqs. (3.3), (3.20) and (3.21), we obtain the required result.
Theorem 3.6. Let $\alpha$ be a spacelike curve in $G_{3}^{1}$ with its pseudo-Galilean trihedron $\{T(s), N(s), B(s)\}$. If the curve $\alpha$ lies on a pseudo-Galilean sphere $S_{ \pm}^{2}$, then it is $N$-constant curve of second kind and the center of a pseudo-Galilean sphere of $\alpha$ at the point $c(s)$ is given by

$$
c(s)=\alpha(s)+m_{1}(s) N(s)+m_{2}(s) B(s) .
$$

Proof. Let $S_{ \pm}^{2}$ be a sphere in $G_{3}^{1}$, then $S_{ \pm}^{2}$ is given by

$$
S_{ \pm}^{2}=\left\{u \in G_{3}^{1}: g(u, u)= \pm r^{2}\right\}
$$

where $r$ is the radius of the pseudo-Galilean sphere and it is a constant. Let $c$ be the center of the pseudo-Galilean sphere, then we have

$$
g(c(s)-\alpha(s), c(s)-\alpha(s))= \pm r^{2}
$$

Differentiating this equation with respect to $s$, we get

$$
\begin{equation*}
g(-T(s), c(s)-\alpha(s))=0 \tag{3.22}
\end{equation*}
$$

more differentiation yields

$$
g\left(-T^{\prime}(s), c(s)-\alpha(s)\right)+g(-T(s),-T(s))=0 .
$$

From Eq. (2.8), we find

$$
\begin{equation*}
-\kappa(s) g(N(s), c(s)-\alpha(s))+1=0 \tag{3.23}
\end{equation*}
$$

and since $c(s)-\alpha(s) \in \operatorname{Sp}\{T(s), N(s), B(s)\}$, then we can write

$$
\begin{equation*}
c(s)-\alpha(s)=m_{o}(s) T(s)+m_{1}(s) N(s)+m_{2}(s) B(s) . \tag{3.24}
\end{equation*}
$$

Now, from Eq. (3.23) and (3.24), we find

$$
\kappa(s) m_{1}(s)+1=0,
$$

it follows that

$$
m_{1}(s)=-\frac{1}{\kappa(s)}
$$

Also, from Eq. (3.22) and (3.24), one can write

$$
g(T(s), c(s)-\alpha(s))=m_{0}(s)
$$

which gives

$$
m_{o}(s)=0,
$$

and then Eq. (3.24) becomes

$$
c(s)-\alpha(s)=m_{1}(s) N(s)+m_{2}(s) B(s) .
$$

Besides, the derivation of Eq. (3.23) leads to

$$
m_{2}(s)=\frac{-m_{1}^{\prime}(s)}{\tau(s)} .
$$

Now, from aforementioned information, we obtain

$$
m_{2}^{2}(s)-m_{1}^{2}(s)= \pm r^{2}=\text { const }
$$

which completes the proof.
Theorem 3.7. Let a be $N$-constant curve of second kind which lies on a pseudo-Galilean sphere $S_{ \pm}^{2}$ with constant radius $r$ in $G_{3}^{1}$. Then

$$
m_{2}^{\prime}(s)-\tau(s) m_{1}(s)=0,
$$

where $m_{2}(s) \neq 0, \tau(s) \neq 0$.
Proof. Let $\alpha$ be a $N$-constant curve in $G_{3}^{1}$, then we have

$$
m_{2}^{2}(s)-m_{1}^{2}(s)= \pm r^{2}
$$

since $r$ is constant, then

$$
m_{2}(s) m_{2}^{\prime}(s)-m_{1}(s) m_{1}^{\prime}(s)=0 .
$$

Substituting $m_{2}(s)=\frac{m_{1}^{\prime}(s)}{\tau(s)}$ in this equation, we get

$$
m_{2}^{\prime}(s)-\tau(s) m_{1}(s)=0 .
$$

Thus, the proof is completed.
Theorem 3.8. Let $\alpha(s)$ be a spacelike curve in $G_{3}^{1}$ with $\kappa(s) \neq 0, \tau(s) \neq 0$. The image of the $N$ constant curve $\alpha$ lies on a pseudo-Galilean sphere $S_{ \pm}^{2}$ if and only if for each $s \in I \subset R$, its curvatures satisfy the following equalities:

$$
\begin{align*}
s+c_{0} & =0 \\
\frac{1}{4} e^{-u(s)}\left(-4 c_{4}+e^{2 u(s)}\right)-\frac{1}{2} e^{u(s)} & =\frac{1}{\kappa(s)} \\
\frac{1}{4} e^{-u(s)}\left(-4 c_{4}+e^{2 u(s)}\right) & =\frac{\kappa^{\prime}(s)}{\kappa^{2}(s) \tau(s)} \tag{3.25}
\end{align*}
$$

where $u(s)=\int \tau(s) d s+c_{5}$ and $c_{0}, c_{4}$ and $c_{5} \in R$.
Proof. By assumption, we have

$$
g(\alpha(s), \alpha(s))=r^{2}
$$

for every $s \in I \subset R$ and $r$ is the radius of the pseudo-Galilean sphere. Differentiating this equation with respect to $s$ gives

$$
\begin{equation*}
g(T(s), \alpha(s))=0 \tag{3.26}
\end{equation*}
$$

Again, differentiation leads to

$$
\begin{equation*}
g(N(s), \alpha(s))=-\frac{1}{\kappa(s)}, \tag{3.27}
\end{equation*}
$$

and also

$$
\begin{equation*}
g(B(s), \alpha(s))=\frac{\kappa^{\prime}(s)}{\kappa^{2}(s) \tau(s)} \tag{3.28}
\end{equation*}
$$

Using Eqs. (3.26)-(3.28) in Eq. (3.19), we obtain the required result: Eq. (3.25).
Conversely, we assume that Eq. (3.25) holds, for each $s \in I \subset R$, then from Eq. (3.19), the position vector of $\alpha$ can be expressed as

$$
\alpha(s)=-\frac{1}{\kappa(s)} N(s)+\frac{\kappa^{\prime}(s)}{\kappa^{2}(s) \tau(s)} B(s),
$$

which satisfies the equation: $g(\alpha(s), \alpha(s))=r^{2}$. It means that the curve $\alpha$ lies on the pseudo-Galilean sphere $S_{ \pm}^{2}$. Hence, the proof is completed.

Theorem 3.9. Let $\alpha$ be a spacelike curve in $G_{3}^{1}$. If $\alpha$ is a circle then $\alpha$ is $N$-constant curve of second kind.

Proof. If $\alpha$ is a circle, then we have

$$
\kappa(s)=\text { const and } \quad \tau(s)=0
$$

Also, from Theorem 3.4, one can write

$$
m_{1}=\frac{1}{c_{3} \kappa}=\text { const., }
$$

$$
m_{2}=\int\left(\frac{-\tau}{c_{3} \kappa}\right) d s=\text { const. }
$$

which leads to

$$
m_{2}^{2}(s)-m_{1}^{2}(s)=\text { const } .
$$

thus, it completes the proof.

## 4. Examples

In this section, we give some examples to illustrate our main results.
Example 4.1. Consider the following spacelike curve $\alpha: I \subset R \rightarrow G_{3}^{1}$, given by

$$
\begin{equation*}
\alpha(s)=\left(s, \frac{s}{6}[2 \sinh (2 \ln s)-\cosh (2 \ln s)], \frac{s}{6}[2 \cosh (2 \ln s)-\sinh (2 \ln s)]\right) . \tag{4.1}
\end{equation*}
$$

Differentiating Eq. (4.1), we get

$$
\begin{equation*}
\alpha^{\prime}(s)=\left(1, \frac{1}{2} \cosh (2 \ln s), \frac{1}{2} \sinh (2 \ln s)\right) . \tag{4.2}
\end{equation*}
$$

Pseudo-Galilean inner product follows that $\left\langle\alpha^{\prime}, \alpha^{\prime}\right\rangle=1$. So the curve is parameterized by the arclength. The tangent vector is

$$
T^{\prime}=\left(0, \frac{1}{s} \sinh (2 \ln s), \frac{1}{s} \cosh (2 \ln s)\right),
$$

by taking the norm of both sides, we have $\kappa(s)=\frac{1}{s}$. Thereafter, we have

$$
N=(0, \sinh (2 \ln s), \cosh (2 \ln s)),
$$

and the binormal vector is

$$
B=(0,-\cosh (2 \ln s),-\sinh (2 \ln s)) .
$$

From Serret-Frenet equations, one can obtain $\tau(s)=\frac{-2}{s}$. Moreover, the curvature functions $m_{i}(s)$ are

$$
m_{0}=s, \quad m_{1}=\frac{s}{c_{3}}, \quad m_{2}=-\Omega s, \Omega=\left(\frac{1+c_{3}}{2 c_{3}}\right)=\text { const. }
$$

So, from Eq. (3.16), we get

$$
\frac{m_{o}^{2}}{m_{2}^{2}-m_{1}^{2}}=\digamma, \quad \digamma=\frac{4\left(c_{3}\right)^{2}}{\left(c_{3}+1\right)^{2}-4}=\text { const. }
$$

Under the above considerations, $\alpha$ is of constant-ratio and the ratio is equal $\digamma$. Also, since

$$
\left\|\alpha^{N}(s)\right\|^{2}=m_{2}^{2}(s)-m_{1}^{2}(s)=\left(\frac{\left(c_{3}+1\right)^{2}-4}{4\left(c_{3}\right)^{2}}\right) s^{2} \neq \text { const. } .
$$

then the curve $\alpha$ is a constant-ratio curve but not $N$-constant curve, see Fig(1a).
Example 4.2. Consider a spacelike curve $\gamma(s)$ in $G_{3}^{1}$ parameterized by

$$
\alpha(s)=\left(s,-a \int\left(\int \sinh \left(\frac{s^{2}}{2}\right) d s\right) d s, a \int\left(\int \cosh \left(\frac{s^{2}}{2}\right) d s\right) d s\right),
$$

where $a \in R$.

Then we have

$$
\begin{aligned}
& \gamma^{\prime}(s)=T(s)=\left(1,-a \int \sinh \left(\frac{s^{2}}{2}\right) d s, a \int \cosh \left(\frac{s^{2}}{2}\right) d s\right), \\
& T^{\prime}(s)=\left(0,-a \sinh \left(\frac{s^{2}}{2}\right), a \cosh \left(\frac{s^{2}}{2}\right)\right) .
\end{aligned}
$$

By a straightforward calculations, we obtain

$$
\begin{aligned}
& N(s)=\left(0,-\sinh \left(\frac{s^{2}}{2}\right), \cosh \left(\frac{s^{2}}{2}\right)\right), \\
& B(s)=\left(0,-\cosh \left(\frac{s^{2}}{2}\right), \sinh \left(\frac{s^{2}}{2}\right)\right),
\end{aligned}
$$

where $\kappa(s)=a=$ const and $\tau(s)=s$.
Since the curve has a constant curvature and non-constant torsion, so it is a Salkowski curve.
From Theorem 3.4, we have the curvature functions:

$$
\begin{aligned}
& m_{1}=\frac{1}{c_{3} \kappa}=\frac{1}{a c_{3}} \\
& m_{2}=\frac{\kappa^{\prime}-\kappa^{3} c_{3} s}{c_{3} \kappa^{2} \tau}=-a, \text { a is constant, }
\end{aligned}
$$

which leads to

$$
m_{2}^{2}(s)-m_{1}^{2}(s)=(-a)^{2}-\left(\frac{1}{a c_{3}}\right)^{2}=\text { const }
$$

It follows that $\gamma$ is $N$-constant curve but not constant-ratio curve, see Fig(1b).


Figure 1. (A) The constant-ratio curve $\alpha,(B)$ the $N$-constant Salkowski curve $\gamma$; $a=2$.

## 5. Conclusion

In the three-dimensional pseudo-Galilean space, spacelike admissible curves of constant-ratio and some special curves such as $T$-constant and N -constant curves have been studied. Furthermore, the spherical images of these curves have been studied. Some interesting results of $N$ - constant curves have been obtained. Finally, as an application for this work, two examples are given and plotted to confirm our main results.
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